

12.5: Perturbative solution of the Liouville equation

Substituting the perturbative form for $f(x, t)$ into the Liouville equation, one obtains

$$\frac{\partial}{\partial t}(f_0(x) + \Delta f(x, t)) + (iL_0 + i\Delta L(t))(f_0(x) + \Delta f(x, t)) = 0$$

Recall $\partial f_0 / \partial t = 0$. Thus, working to linear order in small quantities, one obtains the following equation for $\Delta f(x, t)$:

$$\left(\frac{\partial}{\partial t} + iL_0\right) \Delta f(x, t) = -i\Delta L f_0(x)$$

which is just a first-order inhomogeneous differential equation. This can easily be solved using an integrating factor, and one obtains the result

$$\Delta f(x, t) = -\int_0^t ds e^{-iL_0(t-s)} i\Delta L(s) f_0(x)$$

Note that

$$i\Delta L f_0(x) = iL f_0(x) - iL_0 f_0(x) = iL f_0(x) = \dot{\mathbf{x}} \cdot \nabla_{\mathbf{x}} f_0(x)$$

But, using the chain rule, we have

$$\begin{aligned} \dot{\mathbf{x}} \cdot \nabla_{\mathbf{x}} f_0(x) &= \dot{\mathbf{x}} \cdot \frac{\partial f_0}{\partial \mathbf{H}} \frac{\partial \mathbf{H}}{\partial \mathbf{x}} \\ &= \frac{\partial f_0}{\partial \mathbf{H}} \sum_{i=1}^{3N} \left[\dot{p}_i \frac{\partial \mathbf{H}}{\partial p_i} + \dot{q}_i \frac{\partial \mathbf{H}}{\partial q_i} \right] \\ &= \frac{\partial f_0}{\partial \mathbf{H}} \sum_{i=1}^{3N} \left[\frac{\partial \mathbf{H}}{\partial p_i} \left(-\frac{\partial \mathbf{H}}{\partial q_i} + D_i F_e(t) \right) + \frac{\partial \mathbf{H}}{\partial q_i} \left(\frac{\partial \mathbf{H}}{\partial p_i} + C_i F_e(t) \right) \right] \\ &= \frac{\partial f_0}{\partial \mathbf{H}} \sum_{i=1}^{3N} \left[D_i(x) \frac{\partial \mathbf{H}}{\partial p_i} + C_i(x) \frac{\partial \mathbf{H}}{\partial q_i} \right] F_e(t) \end{aligned}$$

Define

$$j(x) = -\sum_{i=1}^{3N} \left[D_i(x) \frac{\partial \mathbf{H}}{\partial p_i} + C_i(x) \frac{\partial \mathbf{H}}{\partial q_i} \right]$$

which is known as the *dissipative flux*. Thus, for a Cartesian Hamiltonian

$$H = \sum_{i=1}^N \frac{\mathbf{p}_i^2}{2m_i} + U(\mathbf{r}_1, \dots, \mathbf{r}_N)$$

where $\mathbf{F}_i(\mathbf{r}_1, \dots, \mathbf{r}_N) = -\nabla_i U$ is the force on the i th particle, the dissipative flux becomes:

$$j(x) = \sum_{i=1}^N \left[\mathbf{C}_i(x) \cdot \mathbf{F}_i - \mathbf{D}_i(x) \cdot \frac{\mathbf{p}_i}{m_i} \right]$$

In general,

$$\dot{\mathbf{x}} \cdot \nabla_{\mathbf{x}} f_0(x) = -\frac{\partial f_0}{\partial \mathbf{H}} j(x) F_e(t)$$

Now, suppose $f_0(x)$ is a canonical distribution function

$$f_0(H(x)) = \frac{1}{Q(N, V, T)} e^{-\beta H(x)}$$

then

$$\frac{\partial f_0}{\partial H} = -\beta f_0(H)$$

so that

$$\dot{\mathbf{x}} \cdot \nabla_{\mathbf{x}} f_0(\mathbf{x}) = \beta f_0(\mathbf{x}) j(\mathbf{x}) F_e(t)$$

Thus, the solution for $\Delta f(x, t)$ is

$$\Delta f(\mathbf{x}, t) = -\beta \int_0^t ds e^{-iL_0(t-s)} f_0(\mathbf{x}) j(\mathbf{x}) F_e(s)$$

The ensemble average of the observable $A(x)$ now becomes

$$\begin{aligned} \langle A(t) \rangle &= \langle A \rangle_0 - \beta \int d\mathbf{x} A(\mathbf{x}) \int_0^t ds e^{-iL_0(t-s)} f_0(\mathbf{x}) j(\mathbf{x}) F_e(s) \\ &= \langle A \rangle_0 - \beta \int_0^t ds \int d\mathbf{x} A(\mathbf{x}) e^{-iL_0(t-s)} f_0(\mathbf{x}) j(\mathbf{x}) F_e(s) \\ &= \langle A \rangle_0 - \beta \int_0^t ds \int d\mathbf{x} f_0(\mathbf{x}) A(\mathbf{x}) e^{-iL_0(t-s)} j(\mathbf{x}) F_e(s) \end{aligned}$$

Recall that the classical propagator is $exp(iLt)$. Thus the operator appearing in the above expression is a classical propagator of the unperturbed system for propagating backwards in time to $-(t-s)$. An observable $A(x)$ evolves in time according to

$$\begin{aligned} \frac{dA}{dt} &= iLA \\ A(t) &= e^{iLt} A(0) \\ A(-t) &= e^{-iLt} A(0) \end{aligned}$$

Now, if we take the complex conjugate of both sides, we find

$$A^*(t) = A^*(0) e^{-iLt}$$

where now the operator acts to the left on $A^*(0)$. However, since observables are real, we have

$$A(t) = A(0) e^{-iLt}$$

which implies that forward evolution in time can be achieved by acting to the left on an observable with the time reversed classical propagator. Thus, the ensemble average of A becomes

$$\begin{aligned} \langle A(t) \rangle &= \langle A \rangle_0 - \beta \int_0^t ds F_e(s) \int d\mathbf{x}_0 f_0(\mathbf{x}_0) A(\mathbf{x}_{t-s}(\mathbf{x}_0)) j(\mathbf{x}_0) \\ &= \langle A \rangle_0 - \beta \int_0^t ds F_e(s) \langle j(0) A(t-s) \rangle_0 \end{aligned}$$

where the quantity on the last line is an object we have not encountered yet before. It is known as an **equilibrium time correlation function**. An equilibrium time correlation function is an ensemble average over the unperturbed (canonical) ensemble of the product of the dissipative flux at $t=0$ with an observable A evolved to a time $t-s$. Several things are worth noting:

1. The nonequilibrium average $\langle A(t) \rangle$, in the linear response regime, can be expressed solely in terms of equilibrium averages.
2. The propagator used to evolve $A(x)$ to $A(x, t-s)$ is the operator $exp(iL_0(t-s))$, which is the propagator for the unperturbed, Hamiltonian dynamics with $C_i = D_i = 0$. That is, it is just the dynamics determined by H .
3. Since $A(\mathbf{x}, t-s) = A(\mathbf{x}(t-s))$ is a function of the phase space variables evolved to a time $t-s$, we must now specify over which set of phase space variables the integration $\int d\mathbf{x}$ is taken. The choice is actually arbitrary, and for convenience, we choose the initial conditions. Since $x(t)$ is a function of the initial conditions $x(0)$, we can write the time correlation function as

$$\langle j(0) A(t-s) \rangle_0 = \frac{1}{Q} \int d\mathbf{x}_0 e^{-\beta H(\mathbf{x}_0)} j(\mathbf{x}_0) A(\mathbf{x}_{t-s}(\mathbf{x}_0))$$

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