

4.5: Alpha Decay

Many types of heavy atomic nucleus spontaneously decay to produce daughter nuclei via the emission of α -particles (i.e., helium nuclei) of some characteristic energy. This process is known as α -decay. Let us investigate the α -decay of a particular type of atomic nucleus of radius R , charge-number Z , and mass-number A . Such a nucleus thus decays to produce a daughter nucleus of charge-number $Z_1 = Z - 2$ and mass-number $A_1 = A - 4$, and an α -particle of charge-number $Z_2 = 2$ and mass-number $A_2 = 4$. Let the characteristic energy of the α -particle be E . Incidentally, nuclear radii are found to satisfy the empirical formula

$$R = 1.5 \times 10^{-15} A^{1/3} \text{ m} = 2.0 \times 10^{-15} Z_1^{1/3} \text{ m} \quad (4.5.1)$$

for $Z \gg 1$.

In 1928, George Gamow proposed a very successful theory of α -decay, according to which the α -particle moves freely inside the nucleus, and is emitted after tunneling through the potential barrier between itself and the daughter nucleus. In other words, the α -particle, whose energy is E , is trapped in a potential well of radius R by the potential barrier

$$V(r) = \frac{Z_1 Z_2 e^2}{4\pi \epsilon_0 r} \quad (4.5.2)$$

for $r > R$.

Making use of the WKB approximation (and neglecting the fact that r is a radial, rather than a Cartesian, coordinate), the probability of the α -particle tunneling through the barrier is

$$|T|^2 = \exp\left(-\frac{2\sqrt{2m}}{\hbar} \int_{r_1}^{r_2} \sqrt{V(r) - E} dr\right), \quad (4.5.3)$$

where $r_1 = R$ and $r_2 = Z_1 Z_2 e^2 / (4\pi \epsilon_0 E)$. Here, $m = 4 m_p$ is the α -particle mass. The previous expression reduces to

$$|T|^2 = \exp\left(-2\sqrt{2}\beta \int_1^{E_c/E} \left[\frac{1}{y} - \frac{E}{E_c}\right]^{1/2} dy\right), \quad (4.5.4)$$

where

$$\beta = \left(\frac{Z_1 Z_2 e^2 m R}{4\pi \epsilon_0 \hbar^2}\right)^{1/2} = 0.74 Z_1^{2/3} \quad (4.5.5)$$

is a dimensionless constant, and

$$E_c = \frac{Z_1 Z_2 e^2}{4\pi \epsilon_0 R} = 1.44 Z_1^{2/3} \text{ MeV} \quad (4.5.6)$$

is the characteristic energy the α -particle would need in order to escape from the nucleus without tunneling. Of course, $E \ll E_c$. It is easily demonstrated that

$$\int_1^{1/\epsilon} \left(\frac{1}{y} - \epsilon\right)^{1/2} dy \simeq \frac{\pi}{2\sqrt{\epsilon}} - 2 \quad (4.5.7)$$

when $\epsilon \ll 1$. Hence,

$$|T|^2 \simeq \exp\left(-2\sqrt{2}\beta \left[\frac{\pi}{2} \sqrt{\frac{E_c}{E}} - 2\right]\right). \quad (4.5.8)$$

Now, the α -particle moves inside the nucleus with the characteristic velocity $v = \sqrt{2E/m}$. It follows that the particle bounces backward and forward within the nucleus at the frequency $\nu \simeq v/R$, giving

$$\nu \simeq 2 \times 10^{28} \text{ yr}^{-1} \quad (4.5.9)$$

for a 1 MeV α -particle trapped inside a typical heavy nucleus of radius 10^{-14} m. Thus, the α -particle effectively attempts to tunnel through the potential barrier ν times a second. If each of these attempts has a probability $|T|^2$ of succeeding then the probability of

decay per unit time is $\nu |T|^2$. Hence, if there are $N(t) \gg 1$ undecayed nuclei at time t then there are only $N + dN$ at time $t + dt$, where

$$dN = -N \nu |T|^2 dt. \quad (4.5.10)$$

This expression can be integrated to give

$$N(t) = N(0) \exp(-\nu |T|^2 t). \quad (4.5.11)$$

Now, the *half-life*, τ , is defined as the time which must elapse in order for half of the nuclei originally present to decay. It follows from the previous formula that

$$\tau = \frac{\ln 2}{\nu |T|^2}. \quad (4.5.12)$$

Note that the half-life is independent of $N(0)$.

Finally, making use of the previous results, we obtain

$$\log_{10}[\tau(\text{yr})] = -C_1 - C_2 Z_1^{2/3} + C_3 \frac{Z_1}{\sqrt{E(\text{MeV})}}, \quad (4.5.13)$$

where

$$C_1 = 28.5,$$

$$C_2 = 1.83,$$

$$C_3 = 1.73.$$

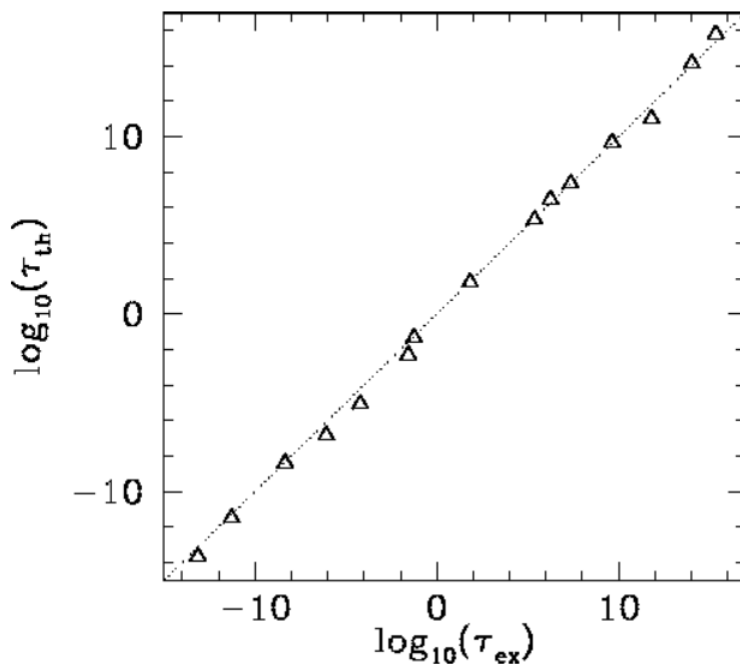


Figure 15: The experimentally determined half-life, τ_{ex} of various atomic nuclei which decay via α emission versus the best-fit theoretical half-life $\log_{10}(\tau_{\text{th}}) = -28.9 - 1.60 Z_1^{2/3} + 1.61 Z_1 / \sqrt{E}$. Both half-lives are measured in years. Here, $Z_1 = Z - 2$. Both half-lives are measured in years. Here, $Z_1 = Z - 2$, where Z is the charge number of the nucleus, and E the characteristic energy of the emitted α -particle in MeV. In order of increasing half-life, the points correspond to the following nuclei: Rn 215, Po 214, Po 216, Po 197, Fm 250, Ac 225, U 230, U 232, U 234, Gd 150, U 236, U 238, Pt 190, Gd 152, Nd 144. Data obtained from IAEA Nuclear Data Centre.

Equation ([e5.64]) is known as the *Geiger-Nuttall formula*, because it was discovered empirically by H. Geiger and J.M. Nuttall in 1911.

The half-life, τ , the daughter charge-number, $Z_1 = Z - 2$, and the α -particle energy, E , for atomic nuclei which undergo α -decay are indeed found to satisfy a relationship of the form ([e5.64](#)). The best fit to the data (see Figure [fal](#)) is obtained using

$$C_1 = 28.9,$$

$$C_2 = 1.60,$$

$$C_3 = 1.61.$$

Note that these values are remarkably similar to those calculated previously.

Contributors and Attributions

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