

## 5.S: Summary

### References

- F. Reif, *Fundamentals of Statistical and Thermal Physics* (McGraw-Hill, 1987) This has been perhaps the most popular undergraduate text since it first appeared in 1967, and with good reason.
- A. H. Carter, *Classical and Statistical Thermodynamics* (Benjamin Cummings, 2000) A very relaxed treatment appropriate for undergraduate physics majors.
- D. V. Schroeder, *An Introduction to Thermal Physics* (Addison-Wesley, 2000) This is the best undergraduate thermodynamics book I've come across, but only 40% of the book treats statistical mechanics.
- C. Kittel, *Elementary Statistical Physics* (Dover, 2004) Remarkably crisp, though dated, this text is organized as a series of brief discussions of key concepts and examples. Published by Dover, so you can't beat the price.
- R. K. Pathria, *Statistical Mechanics* (2<sup>nd</sup> edition, Butterworth-Heinemann, 1996) This popular graduate level text contains many detailed derivations which are helpful for the student.
- M. Plischke and B. Bergersen, *Equilibrium Statistical Physics* (3<sup>rd</sup> edition, World Scientific, 2006) An excellent graduate level text. Less insightful than Kardar but still a good modern treatment of the subject. Good discussion of mean field theory.
- E. M. Lifshitz and L. P. Pitaevskii, *Statistical Physics* (part I, 3<sup>rd</sup> edition, Pergamon, 1980) This is volume 5 in the famous Landau and Lifshitz *Course of Theoretical Physics*. Though dated, it still contains a wealth of information and physical insight.

### Summary

$\hat{H} = \sum_{\alpha} \epsilon_{\alpha} \hat{n}_{\alpha}$ , where  $\epsilon_{\alpha}$  is the energy eigenvalue for the single particle state  $\psi_{\alpha}$  (possibly degenerate), and  $\hat{n}_{\alpha}$  is the number operator. Many-body eigenstates  $|\Psi\rangle$  are labeled by the set of occupancies  $\{n_{\alpha}\}$ , with  $\hat{n}_{\alpha} |\Psi\rangle = n_{\alpha} |\Psi\rangle$ . Thus,  $\hat{H} |\Psi\rangle = E_{\Psi} |\Psi\rangle$ , where  $E_{\Psi} = \sum_{\alpha} n_{\alpha} \epsilon_{\alpha}$ .   
 • **Bosons and fermions**: The allowed values for  $n_{\alpha}$  are  $n_{\alpha} \in \{0, 1, 2, \dots, \infty\}$  for bosons and  $n_{\alpha} \in \{0, 1\}$  for fermions.   
 • **Grand canonical ensemble**: Because of the constraint  $\sum_{\alpha} n_{\alpha} = N$ , the ordinary canonical ensemble is inconvenient. Rather, we use the grand canonical ensemble, in which case

$$\Omega(T, V, \mu) = \pm k_B T \sum_{\alpha} \ln \left( 1 \mp e^{-(\epsilon_{\alpha} - \mu)/k_B T} \right),$$

where the upper sign corresponds to bosons and the lower sign to fermions. The average number of particles occupying the single particle state  $\psi_{\alpha}$  is then

$$\langle \hat{n}_{\alpha} \rangle = \frac{\partial \Omega}{\partial \epsilon_{\alpha}} = \frac{1}{e^{(\epsilon_{\alpha} - \mu)/k_B T} \mp 1}.$$

In the Maxwell-Boltzmann limit,  $\mu \ll -kT$  and  $\langle n_{\alpha} \rangle = z e^{-(\epsilon_{\alpha}/kT)}$ , where  $z = e^{\mu/kT}$  is the fugacity. Note that this low-density limit is common to both bosons and fermions.   
 • **Single particle density of states**: The single particle density of states per unit volume is defined to be

$$g(\epsilon) = \frac{1}{V} \text{Tr} \delta(\epsilon - \hat{h}) = \frac{1}{V} \sum_{\alpha} \delta(\epsilon - \epsilon_{\alpha}),$$

where  $\hat{h}$  is the one-body Hamiltonian. If  $\hat{h}$  is isotropic, then  $\epsilon = \epsilon(k)$ , where  $k = |\mathbf{k}|$  is the magnitude of the wavevector, and

$$g(\epsilon) = \frac{g_d}{(2\pi)^d} \frac{k^{d-1}}{d\epsilon/dk},$$

where  $g_d$  is the degeneracy of each single particle energy state (due to spin, for example).   
 • **Quantum virial expansion**: From  $\Omega = -pV$ , we have

$$n(T, z) = \int_{-\infty}^{\infty} d\varepsilon \frac{g(\varepsilon)}{z^{-1} e^{\varepsilon/k_B T} \mp 1} = \sum_{j=1}^{\infty} (\pm 1)^{j-1} z^j C_j(T)$$

$$\frac{p(T, z)}{k_B T} = \mp \int_{-\infty}^{\infty} d\varepsilon g(\varepsilon) \ln(1 \mp z e^{-\varepsilon/k_B T}) = \sum_{j=1}^{\infty} (\pm 1)^{j-1} \frac{z^j}{j} C_j(T),$$

where

$$C_j(T) = \int_{-\infty}^{\infty} d\varepsilon g(\varepsilon) e^{-j\varepsilon/k_B T}.$$

One now inverts  $n=n(T, z)$  to obtain  $z=z(T, n)$ , then substitutes this into  $p=p(T, z)$  to obtain a series expansion for the equation of state,

$$p(T, n) = nk_B T \left( 1 + B_2(T) n + B_3(T) n^2 + \dots \right).$$

The coefficients  $B_j(T)$  are the virial coefficients. One finds

$$B_2 = \mp \frac{C_2}{2C_1^2}, \quad B_3 = \frac{C_2^2}{C_1^4} - \frac{2C_3}{2C_1^3}.$$

**Photon statistics:** Photons are bosonic excitations whose number is not conserved, hence  $\mu=0$ . The number distribution for photon statistics is then  $n(\varepsilon)=1/(e^{\beta\varepsilon}-1)$ . Examples of particles obeying photon statistics include phonons (lattice vibrations), magnons (spin waves), and of course photons themselves, for which  $\varepsilon(k)=\hbar c k$  with  $g=2$ . The pressure and number density for the photon gas obey  $p(T) = A'_{d,T} T^{d+1}$  and  $n(T)=A'_{d,T} T^d$ , where  $d$  is the dimension of space and  $A_{d,T}$  and  $A'_{d,T}$  are constants. **Blackbody radiation:** The energy density per unit frequency of a three-dimensional blackbody is given by

$$\varepsilon(\nu, T) = \frac{8\pi h}{c^3} \cdot \frac{\nu^3}{e^{h\nu/k_B T} - 1}.$$

The total power emitted per unit area of a blackbody is  $\{dP\over dA\}=\sigma T^4$ , where  $\sigma=\pi^2 k_{ssr}^4/60\hbar^3 c^2=5.67\times 10^{-8}\text{W}/\text{m}^2\text{K}^4$  is Stefan's constant. **Ideal Bose gas:** For Bose systems, we must have  $\varepsilon_{\alpha} > \mu$  for all single particle states. The number density is

$$n(T, \mu) = \int_{-\infty}^{\infty} d\varepsilon \frac{g(\varepsilon)}{e^{\beta(\varepsilon-\mu)} - 1}.$$

This is an increasing function of  $\mu$  and an increasing function of  $T$ . For fixed  $T$ , the largest value  $n(T, \mu)$  can attain is  $n(T, \varepsilon_{ns_0})$ , where  $\varepsilon_{ns_0}$  is the lowest possible single particle energy, for which  $g(\varepsilon)=0$  for  $\varepsilon < \varepsilon_{ns_0}$ . If  $n_{ns_{Rc}}(T) \equiv n(T, \varepsilon_{ns_0}) < \infty$ , this establishes a critical density above which there is Bose condensation into the energy  $\varepsilon_{ns_0}$  state. Conversely, for a given density  $n$  there is a critical temperature  $T_{ns_{Rc}}(n)$  such that  $\varepsilon_{ns_0}$  is finite for  $T < T_{ns_{Rc}}$ ,  $n(T, \mu)$  is given by the integral formula above, with  $\varepsilon_{ns_0}=0$ . For a ballistic dispersion  $\varepsilon(k)=\hbar^2 k^2/2m$ , one finds  $n_{ns_{Rc}}(T)^{d/2} = Sg \zeta(d/2)$ ,  $k_B T_{ns_{Rc}} = \{2\pi\hbar^2/m\} \left( n_{ns_{Rc}} Sg \zeta(d/2) \right)^{2/d}$ . For  $T < T_{ns_{Rc}}(n)$ , one has  $n = Sg \text{Li}_{\{d/2+1\}}(z)$ ,  $\lambda T_{ns_{Rc}}^{d/2} = \{2\pi\hbar^2/m\}^{d/2} \text{Li}_{\{d/2+1\}}(z)$ , where

$$Li_q(z) \equiv \sum_{n=1}^{\infty} \frac{z^n}{n^q}.$$

**Ideal Fermi gas:** The Fermi distribution is  $n(\varepsilon)=f(\varepsilon-\mu)=1/(e^{\beta(\varepsilon-\mu)}+1)$ . At  $T=0$ , this is a step function:  $n(\varepsilon)=\Theta(\mu-\varepsilon)$ , and  $n=\int_{-\infty}^{\mu} d\varepsilon g(\varepsilon)$ . The chemical potential at  $T=0$  is called the Fermi energy:  $\mu(T=0, n)=\varepsilon_F(n)$ . If the dispersion is  $\varepsilon(k)$ , the locus of  $k$  values satisfying  $\varepsilon(k)=\varepsilon_F$  is called the Fermi surface. For an isotropic and monotonic dispersion  $\varepsilon(k)$ , the Fermi surface

is a sphere of radius  $k_F$ , the Fermi wavevector. For isotropic three-dimensional systems,  $k_F = (6\pi^2 n)^{1/3}$ .  
 • Sommerfeld expansion: Let  $\phi(\mu) = \frac{d\Phi}{d\mu}$ . Then

$$\int_{-\infty}^{\infty} d\varepsilon f(\varepsilon - \mu) \phi(\varepsilon) = \pi D \csc(\pi D) \Phi(\mu)$$

$$= \left\{ 1 + \frac{\pi^2}{6} (k_B T)^2 \frac{d^2}{d\mu^2} + \frac{7\pi^4}{360} (k_B T)^4 \frac{d^4}{d\mu^4} + \dots \right\} \Phi(\mu),$$

where  $D = k_B T / \mu$ . One then finds, for example,  $\chi_V = \gamma V T$  with  $\gamma = \frac{1}{3} \pi^2 k_B^2 g(\mu_F)$ .

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