

9.9: Summary of Relations Between f , A and S

In this section I use ϖf to mean either $\varpi_1 f_{12}$ or $\varpi_2 f_{21}$, since these are equal; likewise I use ϖB to mean either $\varpi_1 B_{12}$ or $\varpi_2 B_{21}$. The Einstein A coefficient is used exclusively in connection with emission spectroscopy. The B coefficient is defined here in terms of radiation energy density per unit wavelength interval; that is, it is the B^a of section 9.4. The relations between the possible definitions of B are given in equations 9.4.1-4.

The following relations for electric dipole radiation may be useful. In these, ϵ_0 is the “rationalized” definition of free space permittivity, and the formulas are suitable for use with SI units.

$$\varpi_2 A_{21} = \frac{8\pi hc}{\lambda^5} \varpi B = \frac{2\pi e^2}{\epsilon_0 mc \lambda^2} \varpi f = \frac{16\pi^3}{3h\epsilon_0 \lambda^3} S; \quad (9.9.1)$$

$$\varpi B = \frac{e^2 \lambda^3}{4h\epsilon_0 mc^2} \varpi f = \frac{2\pi^2 \lambda^2}{3h^2 \epsilon_0 c} S = \frac{\lambda^5}{8\pi hc} \varpi_2 A_{21}; \quad (9.9.2)$$

$$\varpi f = \frac{8\pi^2 mc}{3he^2 \lambda} S = \frac{\epsilon_0 mc \lambda^2}{2\pi e^2} \varpi_2 A_{21} = \frac{4h\epsilon_0 mc^2}{e^2 \lambda^3} \varpi B; \quad (9.9.3)$$

$$S = \frac{3h\epsilon_0 \lambda^3}{16\pi^3} \varpi_2 A_{21} = \frac{3h^2 \epsilon_0 c}{2\pi^2 \lambda^2} \varpi B = \frac{3he^2 \lambda}{8\pi^2 mc} \varpi f. \quad (9.9.4)$$

For electric quadrupole radiation:

$$\varpi_2 A_{21} = \frac{8\pi^5}{5\epsilon_0 h \lambda^5} S. \quad (9.9.5)$$

For magnetic dipole radiation:

$$\varpi_2 A_{21} = \frac{16\pi^3 \mu_0}{3h \lambda^3} S, \quad (9.9.6)$$

in which μ_0 is the free space permeability.

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