

## 13.6: Distribution of Molecular Speeds

### Learning Objectives

By the end of this section, you will be able to:

- Describe the distribution of molecular speeds in an ideal gas
- Find the average and most probable molecular speeds in an ideal gas

Particles in an ideal gas all travel at relatively high speeds, but they do not travel at the same speed. The rms speed is one kind of average, but many particles move faster and many move slower. The actual distribution of speeds has several interesting implications for other areas of physics, as we will see in later chapters.

### 13.6.1 The Maxwell-Boltzmann Distribution

The motion of molecules in a gas is random in magnitude and direction for individual molecules, but a gas of many molecules has a predictable distribution of molecular speeds. This predictable distribution of molecular speeds is known as the **Maxwell-Boltzmann distribution**, after its originators, who calculated it based on kinetic theory, and it has since been confirmed experimentally (Figure 13.6.1).

To understand this figure, we must define a distribution function of molecular speeds, since with a finite number of molecules, the probability that a molecule will have exactly a given speed is 0.

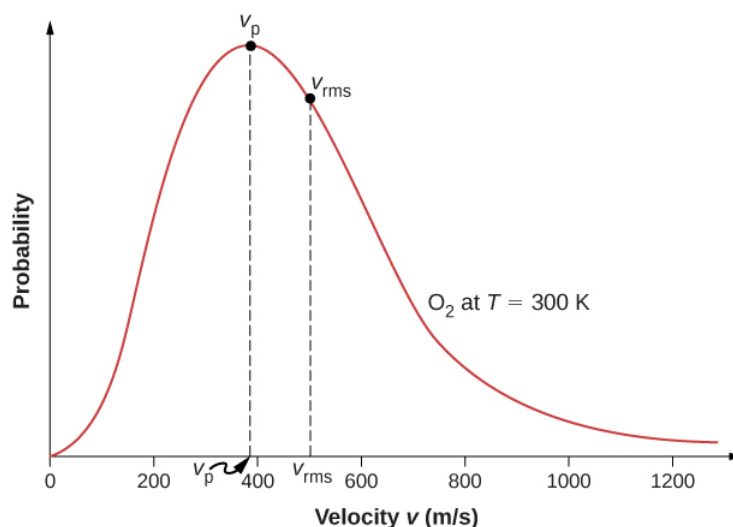


Figure 13.6.1: The Maxwell-Boltzmann distribution of molecular speeds in an ideal gas. The most likely speed  $v_p$  is less than the rms speed  $v_{rms}$ . Although very high speeds are possible, only a tiny fraction of the molecules have speeds that are an order of magnitude greater than  $v_{rms}$ .

We define the distribution function  $f(v)$  by saying that the expected number  $N(v_1, v_2)$  of particles with speeds between  $v_1$  and  $v_2$  is given by

$$N(v_1, v_2) = N \int_{v_1}^{v_2} f(v) dv.$$

[Since  $N$  is dimensionless, the unit of  $f(v)$  is seconds per meter.] We can write this equation conveniently in differential form:

$$dN = N f(v) dv. \quad (13.6.1)$$

In this form, we can understand the equation as saying that the number of molecules with speeds between  $v$  and  $v + dv$  is the total number of molecules in the sample times  $f(v)$  times  $dv$ . That is, the probability that a molecule's speed is between  $v$  and  $v + dv$  is  $f(v)dv$ .

We can now quote Maxwell's result, although the proof is beyond our scope.

## Maxwell-Boltzmann Distribution of Speeds

The distribution function for speeds of particles in an ideal gas at temperature  $T$  is

$$f(v) = \frac{4}{\sqrt{\pi}} \left( \frac{m}{2k_B T} \right)^{3/2} v^2 e^{-mv^2/2k_B T}. \quad (13.6.2)$$

The factors before the  $v^2$  are a normalization constant; they make sure that  $N(0, \infty) = N$  by making sure that  $\int_0^{\infty} f(v) dv = 1$ . Let's focus on the dependence on  $v$ . The factor of  $v^2$  means that  $f(0) = 0$  and for small  $v$ , the curve looks like a parabola. The factor of  $e^{-mv^2/2k_B T}$  means that  $\lim_{v \rightarrow \infty} f(v) = 0$  and the graph has an exponential tail, which indicates that a few molecules may move at several times the rms speed. The interaction of these factors gives the function the single-peaked shape shown in the figure.

### Example 13.6.1: Calculating the Ratio of Numbers of Molecules Near Given Speeds

In a sample of nitrogen  $N_2$  with a molar mass of 28.0 g/mol at a temperature of  $273^\circ C$  find the ratio of the number of molecules with a speed very close to 300 m/s to the number with a speed very close to 100 m/s.

#### Strategy

Since we're looking at a small range, we can approximate the number of molecules near 100 m/s as  $dN_{100} = f(100 \text{ m/s}) dv$ . Then the ratio we want is

$$\frac{dN_{300}}{dN_{100}} = \frac{f(300 \text{ m/s}) dv}{f(100 \text{ m/s}) dv} = \frac{f(300 \text{ m/s})}{f(100 \text{ m/s})}.$$

All we have to do is take the ratio of the two  $f$  values.

#### Solution

1. Identify the knowns and convert to SI units if necessary.

$$T = 300 \text{ K}, k_B = 1.38 \times 10^{-23} \text{ J/K} \quad (13.6.3)$$

$$M = 0.0280 \text{ kg/mol so } m = 4.65 \times 10^{-26} \text{ kg} \quad (13.6.4)$$

2. Substitute the values and solve.

$$\begin{aligned} \frac{f(300 \text{ m/s})}{f(100 \text{ m/s})} &= \frac{\frac{4}{\sqrt{\pi}} \left( \frac{m}{2k_B T} \right)^{3/2} (300 \text{ m/s})^2 \exp[-m(300 \text{ m/s})^2/2k_B T]}{\frac{4}{\sqrt{\pi}} \left( \frac{m}{2k_B T} \right)^{3/2} (100 \text{ m/s})^2 \exp[-m(100 \text{ m/s})^2/2k_B T]} \\ &= \frac{(300 \text{ m/s})^2 \exp[-(4.65 \times 10^{-26} \text{ kg})(300 \text{ m/s})^2/2(1.38 \times 10^{-23} \text{ J/K})(300 \text{ K})]}{(100 \text{ m/s})^2 \exp[-(4.65 \times 10^{-26} \text{ kg})(100 \text{ m/s})^2/2(1.38 \times 10^{-23} \text{ J/K})(300 \text{ K})]} \\ &= 3^2 \exp \left[ -\frac{(4.65 \times 10^{-26} \text{ kg})[(300 \text{ m/s})^2 - (100 \text{ m/s})^2]}{2(1.38 \times 10^{-23} \text{ J/K})(300 \text{ K})} \right] = 5.74 \end{aligned}$$

Figure 13.6.2 shows that the curve is shifted to higher speeds at higher temperatures, with a broader range of speeds.

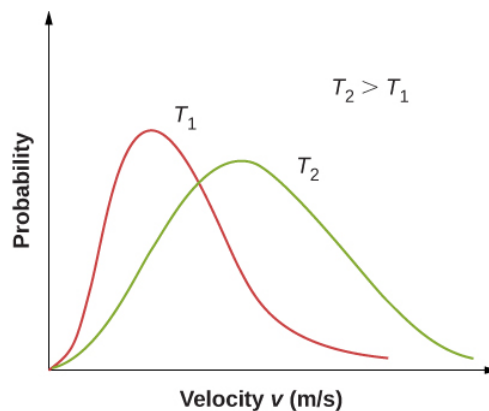


Figure 13.6.2: The Maxwell-Boltzmann distribution is shifted to higher speeds and broadened at higher temperatures.

With only a relatively small number of molecules, the distribution of speeds fluctuates around the Maxwell-Boltzmann distribution. However, you can view this simulation to see the essential features that more massive molecules move slower and have a narrower distribution. Use the set-up “2 Gases, Random Speeds”. Note the display at the bottom comparing histograms of the speed distributions with the theoretical curves.

We can use a probability distribution to calculate average values by multiplying the distribution function by the quantity to be averaged and integrating the product over all possible speeds. (This is analogous to calculating averages of discrete distributions, where you multiply each value by the number of times it occurs, add the results, and divide by the number of values. The integral is analogous to the first two steps, and the normalization is analogous to dividing by the number of values.) Thus the average velocity is

$$\bar{v} = \int_0^{\infty} v f(v) dv = \sqrt{\frac{8}{\pi} \frac{k_B T}{m}} = \sqrt{\frac{8}{\pi} \frac{RT}{M}}. \quad (13.6.5)$$

Similarly,

$$v_{rms} = \sqrt{\bar{v}^2} = \sqrt{\int_0^{\infty} v^2 f(v) dv} = \sqrt{\frac{3k_B T}{m}} = \sqrt{\frac{3RT}{M}} \quad (13.6.6)$$

as in [Pressure, Temperature, and RMS Speed](#). The **most probable speed**, also called the **peak speed**  $v_p$  is the speed at the peak of the velocity distribution. (In statistics it would be called the mode.) It is less than the rms speed  $v_{rms}$ . The most probable speed can be calculated by the more familiar method of setting the derivative of the distribution function, with respect to  $v$ , equal to 0. The result is

$$v_p = \sqrt{\frac{2k_B T}{m}} = \sqrt{\frac{2RT}{M}}, \quad (13.6.7)$$

which is less than  $v_{rms}$ . In fact, the rms speed is greater than both the most probable speed and the average speed.

The peak speed provides a sometimes more convenient way to write the Maxwell-Boltzmann distribution function:

$$f(v) = \frac{4v^2}{\sqrt{\pi}v_p^3} e^{-v^2/v_p^2} \quad (13.6.8)$$

In the factor  $e^{-mv^2/2k_B T}$ , it is easy to recognize the translational kinetic energy. Thus, that expression is equal to  $e^{-K/k_B T}$ . The distribution  $f(\mathbf{v})$  can be transformed into a kinetic energy distribution by requiring that  $f(K)dK = f(v)dv$ . Boltzmann showed that the resulting formula is much more generally applicable if we replace the kinetic energy of translation with the total mechanical energy  $E$ . Boltzmann's result is

$$f(E) = \frac{2}{\sqrt{\pi}} (k_B T)^{-3/2} \sqrt{E} e^{-E/k_B T} = \frac{2}{\sqrt{\pi} (k_B T)^{3/2}} \frac{\sqrt{E}}{e^{E/k_B T}}. \quad (13.6.9)$$

The first part of this equation, with the negative exponential, is the usual way to write it. We give the second part only to remark that  $e^{E/k_B T}$  in the denominator is ubiquitous in quantum as well as classical statistical mechanics.

### Problem-Solving Strategy: Speed Distribution

- **Step 1.** Examine the situation to determine that it relates to the distribution of molecular speeds.
- **Step 2.** Make a list of what quantities are given or can be inferred from the problem as stated (identify the known quantities).
- **Step 3.** Identify exactly what needs to be determined in the problem (identify the unknown quantities). A written list is useful.
- **Step 4.** Convert known values into proper SI units (K for temperature, Pa for pressure,  $m^3$  for volume, molecules for  $N$ , and moles for  $n$ ). In many cases, though, using  $R$  and the molar mass will be more convenient than using  $k_B$  and the molecular mass.
- **Step 5.** Determine whether you need the distribution function for velocity or the one for energy, and whether you are using a formula for one of the characteristic speeds (average, most probably, or rms), finding a ratio of values of the distribution function, or approximating an integral.
- **Step 6.** Solve the appropriate equation for the ideal gas law for the quantity to be determined (the unknown quantity). Note that if you are taking a ratio of values of the distribution function, the normalization factors divide out. Or if approximating an integral, use the method asked for in the problem.
- **Step 7.** Substitute the known quantities, along with their units, into the appropriate equation and obtain numerical solutions complete with units.

We can now gain a qualitative understanding of a puzzle about the composition of Earth's atmosphere. Hydrogen is by far the most common element in the universe, and helium is by far the second-most common. Moreover, helium is constantly produced on Earth by radioactive decay. Why are those elements so rare in our atmosphere? The answer is that gas molecules that reach speeds above Earth's escape velocity, about 11 km/s, can escape from the atmosphere into space. Because of the lower mass of hydrogen and helium molecules, they move at higher speeds than other gas molecules, such as nitrogen and oxygen. Only a few exceed escape velocity, but far fewer heavier molecules do. Thus, over the billions of years that Earth has existed, far more hydrogen and helium molecules have escaped from the atmosphere than other molecules, and hardly any of either is now present.

We can also now take another look at evaporative cooling, which we discussed in the chapter on temperature and heat. Liquids, like gases, have a distribution of molecular energies. The highest-energy molecules are those that can escape from the intermolecular attractions of the liquid. Thus, when some liquid evaporates, the molecules left behind have a lower average energy, and the liquid has a lower temperature.

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